



Errors Due to the Reflectivity of Calibration Targets

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I. INTRODUCTION

- NIST microwave radiometry effort
 - 30+ years experience in noise and antenna metrology at NIST; recently began doing remote-sensing radiometry, combining the two.
- Developing microwave radiometry cal stds:
 - E.g., Randa *et al.*, “Standard Radiometers and Targets for Microwave Remote Sensing,” IGARSS-04
- Related issue is the (non-ideal) reflectivity of calibration targets. This work describes:
 - Causes
 - Expressions (approximate) for T_B error introduced
 - Measurement examples



- Calibration targets close to the sensing antenna:
 - linear radiometers need \geq two standards for calibration.
 - satellites: cold sky, if possible (far-field)
 - otherwise: hot & cold targets (near-field)
 - Scene is always far-field
- Near-field targets introduce two general types of error in a total-power radiometer:
 - Antenna+target affects antenna pattern, directivity (ignore)
 - ΔT at antenna output due to non-ideal target (this work):
 - Difference in M (mismatch factor) for target, scene
 - Difference in system F and G_{av} “ “ “



OUTLINE

- Theoretical Framework
 - Radiometer equation (common approximation)
 - Modifications for ΔT effects
 - T_B uncertainty estimates
- Measurements
 - AIMR radiometer, antenna & target (37 GHz)
 - NASA target & NOAA antenna (54 GHz)
- Numerical estimates
- Summary

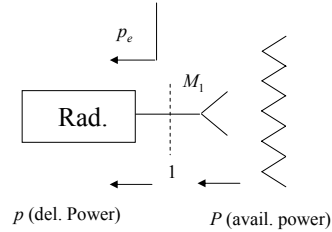


II. THEORETICAL FRAMEWORK

Radiometer equation:

$$p_1 = M_1 P_1 + p_{e,1}$$

$$M_1 = \frac{(1 - |\Gamma_{ant}|^2)(1 - |\Gamma_r|^2)}{|1 - \Gamma_{ant} \Gamma_r|^2}$$



- View cal targets h and c , unknown scene x .
- If M_1 and $p_{e,1}$ are the same for all three cases (h , c , and x) then

$$(T_x - T_c)_0 = \frac{(p_x - p_c)}{(p_h - p_c)} (T_h - T_c)$$

...but what if M_1 and $p_{e,1}$ are *not* the same for all three cases?



Modified Radiometer Equation

- Including the effect of differences in M_1 and $p_{e,1}$ for the three cases:

$$T_x - T_c = (T_x - T_c)_0 (1 + \delta_1) + \Delta_2 + \Delta_3$$

Mismatch changes

F_N and G_{av} changes

[Details described in upcoming TGARSS paper]



Useful approximations for δ_1 , Δ_2 and Δ_3 :

- Assume antenna $\Gamma_h = \Gamma_c$
- Γ_∞ is refl. coeff. of ant. looking at distant scene
- Γ_r is refl. coeff. of radiometer at plane 1

$$\begin{aligned}\delta_1 &\approx 2 \operatorname{Re}[(\Gamma_r - \Gamma_\infty) \Delta \Gamma], \\ \Delta_2 &\approx 2 T_c \operatorname{Re}[(\Gamma_r - \Gamma_\infty) \Delta \Gamma] = \delta_1 T_c, \\ \Delta_3 &\approx 2 X_1 \operatorname{Re}(\Gamma_\infty \Delta \Gamma) + 2 \operatorname{Re}(X_{12} \Delta \Gamma),\end{aligned}$$

where X_1 and X_{12} are noise parameters of the radiometer.
So, we need to know or estimate Γ_r , Γ_∞ , $\Delta \Gamma$, X_1 , and X_{12} .

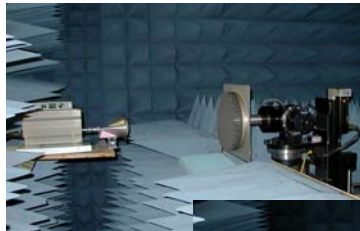
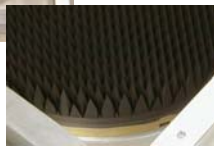
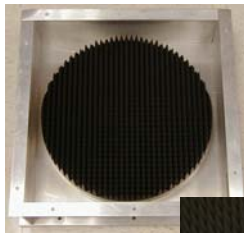
Measure with ANA



III. MEASUREMENTS

- Measured Γ_c , Γ_∞ (thus $\Delta \Gamma$) with ANA for several combinations of antenna and target.

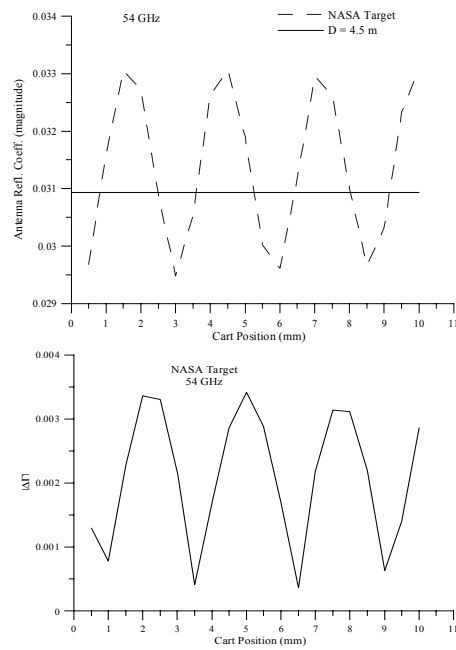
NASA target and NOAA antenna





*NASA target,
NOAA antenna*
 Γ_c, Γ_∞

$$\Delta\Gamma = \Gamma_c - \Gamma_\infty$$

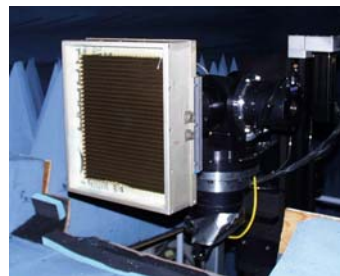
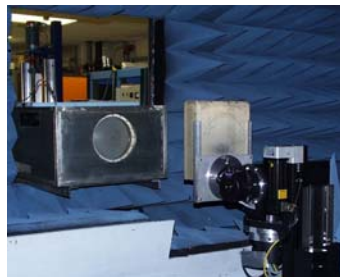


**NIST
NOISE**



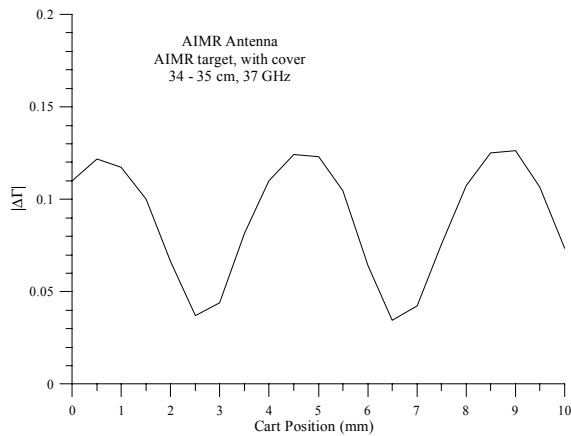
AIMR Antenna & Target

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*AIMR target
and antenna*
 $\Delta\Gamma = \Gamma_c - \Gamma_\infty$



Receiver Noise Parameters

- Use Meys' method to measure X_1 and $|X_{12}|$ at AIMR receiver input at 37 GHz:

- $X_1 \approx 223$ K
- $|X_{12}| \approx 37.6$ K

- For a total-power radiometer with an input isolator, $X_1 \approx T_I$, $X_{12} \approx -T_I S_{11}^I$,

$$\mathcal{A}_2 + \mathcal{A}_3 = 0$$

→ Only remaining error is δ_1



IV. NUMERICAL ESTIMATES



- Total error introduced by using simple form of radiometer equation depends on Γ_r ; assume it's 0.

$$\Delta_{tot}^{(0)} \approx 2(X_1 - T_{x,0}) \text{Re}(\Gamma_\infty \Delta\Gamma) + 2 \text{Re}(X_{12} \Delta\Gamma)$$

For actual rad., use $(\Gamma_r - \Gamma_\infty)$

- Standard uncertainty u given by

$$u_{tot}^{(0)} = \sqrt{\langle (\Delta_{tot}^{(0)})^2 \rangle}$$

RMS over reasonable values of unknown parameters:
ant → target dist. $\angle X_{12} - \angle \Delta\Gamma$

$$= 2 \left\{ (X_1 - T_{x,0})^2 \langle (\text{Re}(\Gamma_\infty \Delta\Gamma))^2 \rangle + \frac{1}{2} |X_{12}|^2 \langle |\Delta\Gamma|^2 \rangle \right\}^{1/2}$$



Uncertainties



- AIMR antenna & target, $T_{x,0}$ from 200 K to 300 K:

$$u_{tot}^{(0)} \approx \sqrt{2} |X_{12}| |\Delta\Gamma|_{RMS} \approx 5.2 K$$

- Prior AIMR cal checks show agreement to within ~2 K
 - $|X_{12}|$ may be overestimated due to meas. time span
 - Spare feedhorn w/o reflector may differ from actual components
 - RMS value is an average; actual instrument is just one position
- Add input isolator with $|S_{11}|=0.025$; for $|T_{x,0} - T_a| \leq 50 K$

$$u_{tot}^{(0)} \approx 1 K$$



Uncertainties

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- NOAA ant., NASA target, $T_{x,0}$ from 200 K-300 K:

$$u_{tot}^{(0)} \approx 0.0033 |X_{12}|$$

- $|X_{12}|$ could be 100 K or more, so uncertainty could be ≥ 0.3 K
 - Significant for some radiometers to be deployed in the next decade
- With an input isolator:

$$u_{tot}^{(0)} \approx 0.95 K \times |S_{11}^I|$$

$$u_{tot}^{(0)} \leq 0.1 K$$



V. SUMMARY

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- Considered error arising from difference in $\Gamma_{ant.}$ when viewing distant scene and nearby cal target
- Developed expressions for approximate T_B error; performed measurements enabling us to estimate u for representative cases
- For total-power radiometers w/o isolators, u can be several kelvins (tenths in good cases)
 - Sensitive to $\Gamma_{ant.}$, Γ_r , target reflectivity, rcvr X 's, and **antenna-target spacing**
- Should measure (or estimate) Γ 's and X 's to estimate uncertainties
- Could correct for these effects w/full rad eq'n



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